## STEREO Observations of Langmuir Waves and Ion Events in the Terrestrial Foreshock

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# Outline

Introduction

Early STEREO Orbit

Electron Foreshock

Ion Foreshock

Conclusions and Future Work

Introduction

**Early STEREO Orbit** 

**Electron Foreshock** 

**Ion Foreshock** 

## The Foreshock

#### Introduction

#### The Foreshock

Instruments

Early STEREO Orbit

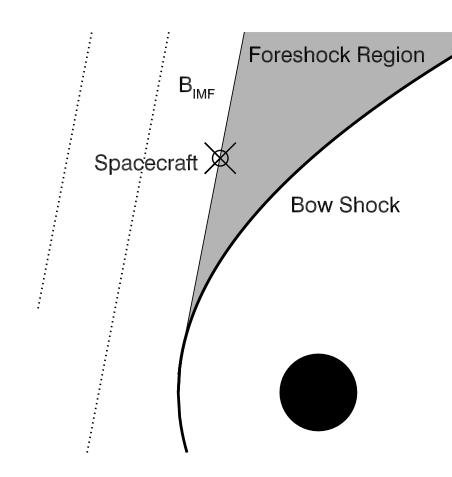
Electron Foreshock

Ion Foreshock

Conclusions and Future Work

- The foreshock is the region upstream of Earth's bow shock that is magnetically connected to the shock.
- Particles in the interplanetary medium travel along field lines and can be mirrored by the bow shock.
- The reflected beams cause bump-on-tail velocity distributions and generate Langmuir wave emission.

(Filbert and Kellogg 1979 and many others)



## Instruments

Introduction

The Foreshock

#### Instruments

Early STEREO Orbit

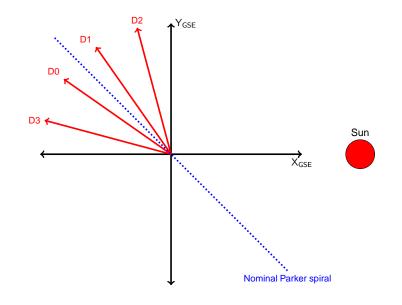
Electron Foreshock

Ion Foreshock

Conclusions and Future Work

### STE:

- 2-100keV
- Solid state detector
- Electrons, ions, neutrals
- 4 look directions, downstream and upstream (STE-U is not working)



### SWEA:

- 8-2000 eV
- Electrostatic detector
- Nearly  $4\pi$  angular resolution

# Late 2006: Petal Orbit

#### Introduction

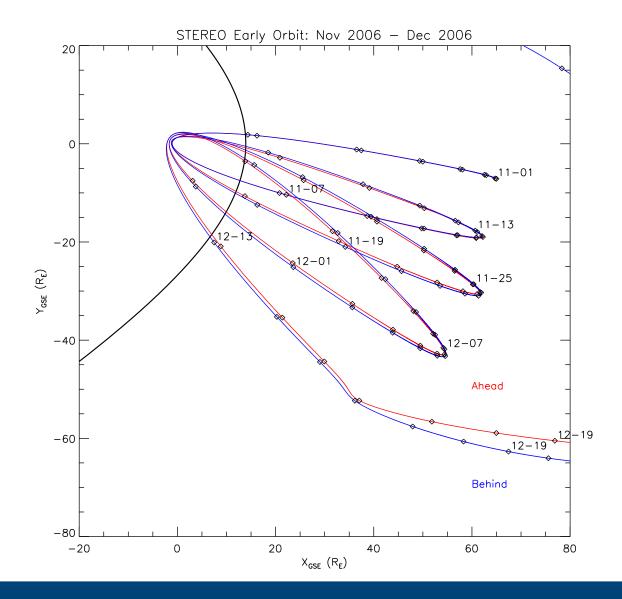
Early STEREO Orbit

Late 2006: Petal Orbit

Early 2007: STEREO/Ahead in the Foreshock

Electron Foreshock

Ion Foreshock



# Early 2007: STEREO/Ahead in the Foreshock

#### Introduction

Early STEREO Orbit Late 2006: Petal

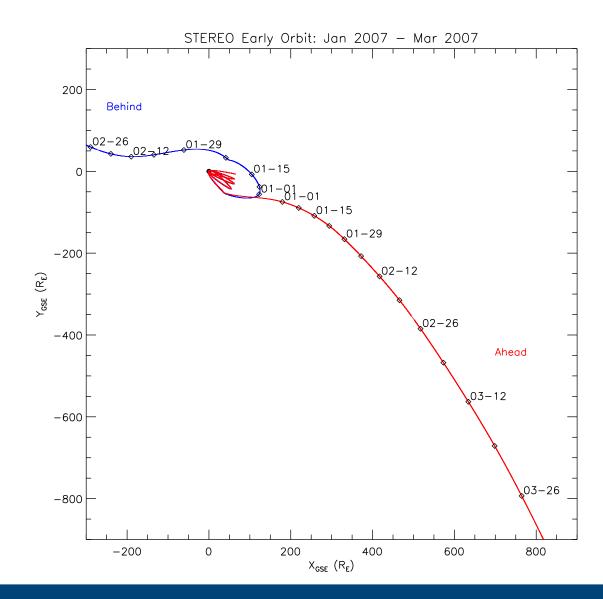
Orbit

Early 2007: STEREO/Ahead in

the Foreshock

Electron Foreshock

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#### Introduction

#### Early STEREO Orbit

#### Electron Foreshock

Electron Event

Electron spectra

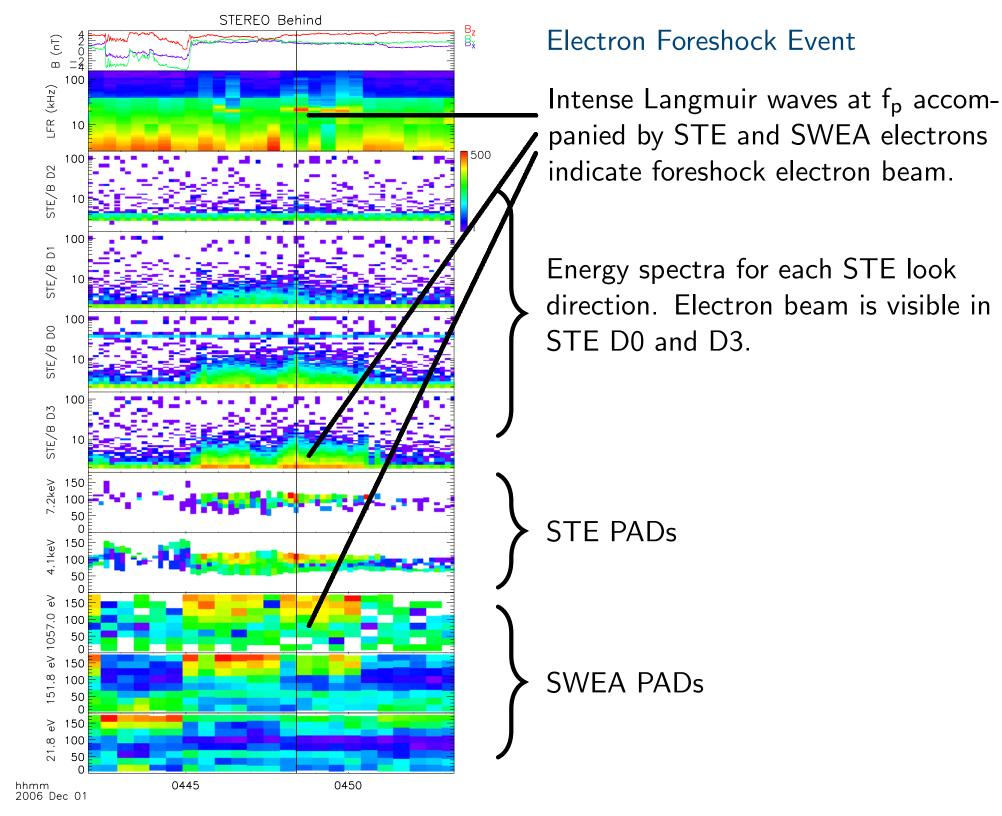
Distribution of

electron events

#### Ion Foreshock

Conclusions and Future Work

# **Electron Foreshock**



## **Electron spectra**

Introduction

Early STEREO Orbit

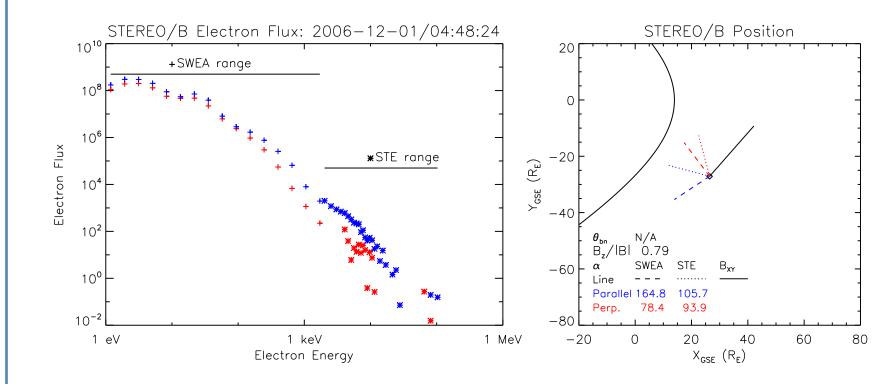
Electron Foreshock

Electron Event

Electron spectra

Distribution of electron events

Ion Foreshock



- Blue = Parallel, Red = Perpendicular.
- Electron beam present in both SWEA and STE.
- Pitch angle coverage limited by fixed look directions of STE.

## Distribution of electron events

#### Introduction

Early STEREO Orbit

Electron Foreshock

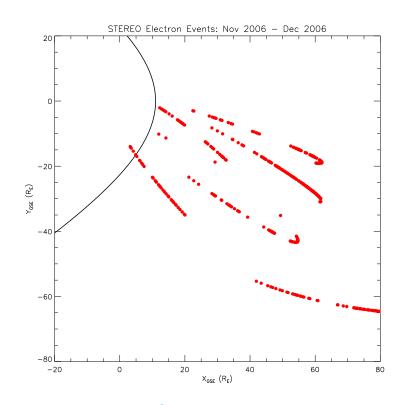
Electron Event

Electron spectra

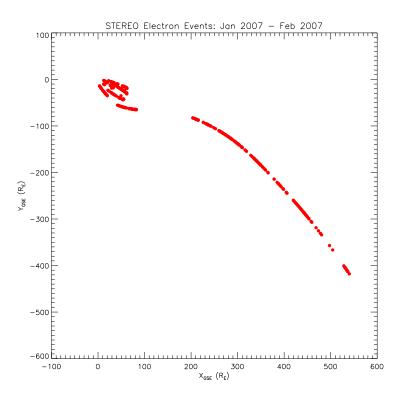
Distribution of electron events

Ion Foreshock

Conclusions and Future Work



Hundreds of events, no immediately obvious organization by spacecraft position.



Electrons continue to be observed months into 2007. (The blank space is due to incomplete cataloguing, not a data gap or a dearth of events.)

#### Introduction

Early STEREO Orbit

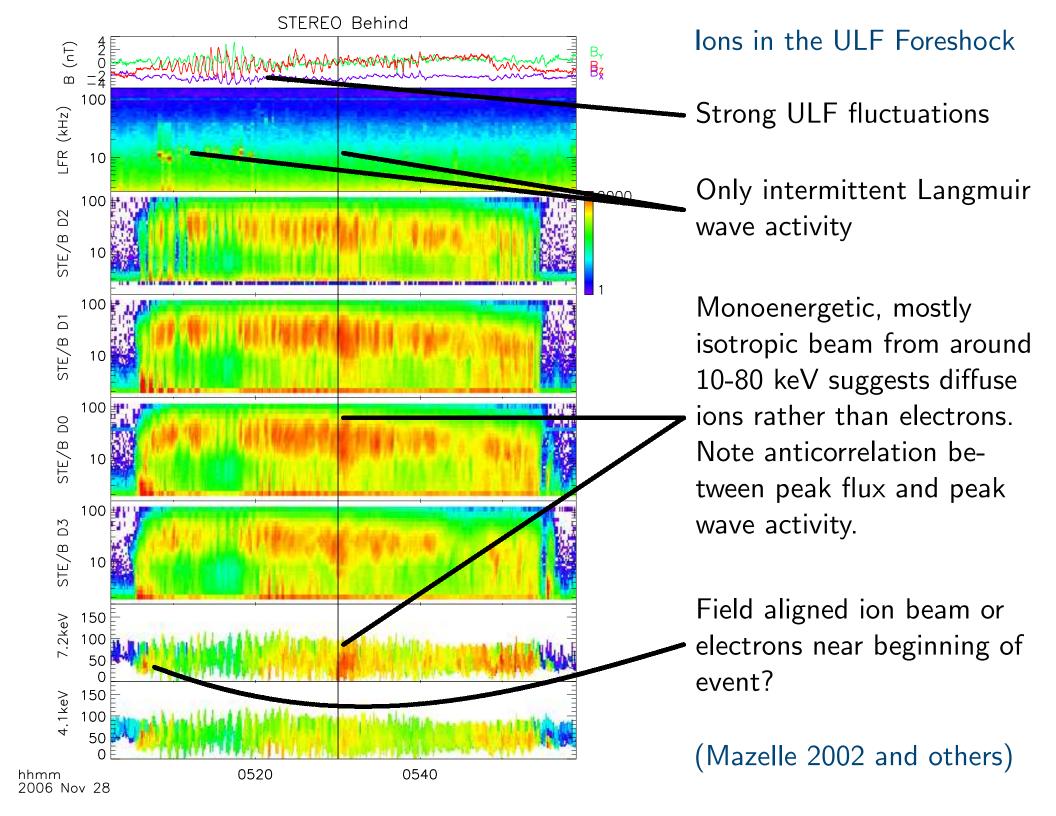
Electron Foreshock

#### Ion Foreshock

Ion Foreshock
Ion spectra
Foreshock event
with fine structure
Distribution of ion
events

Conclusions and Future Work

# **Ion Foreshock**



# lon spectra

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Electron Foreshock

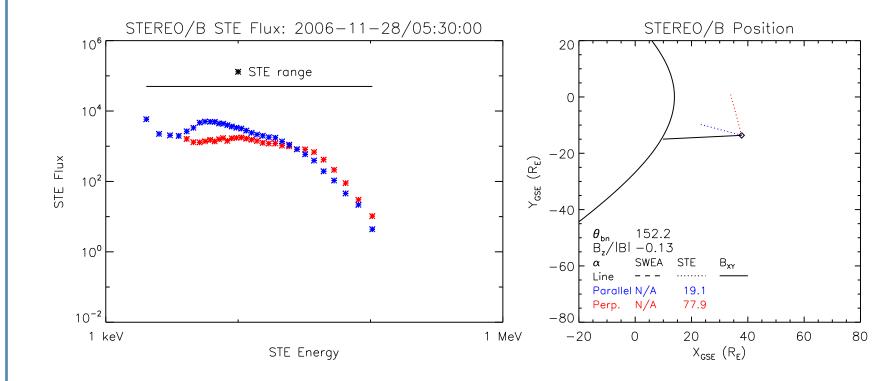
Ion Foreshock

Ion Foreshock

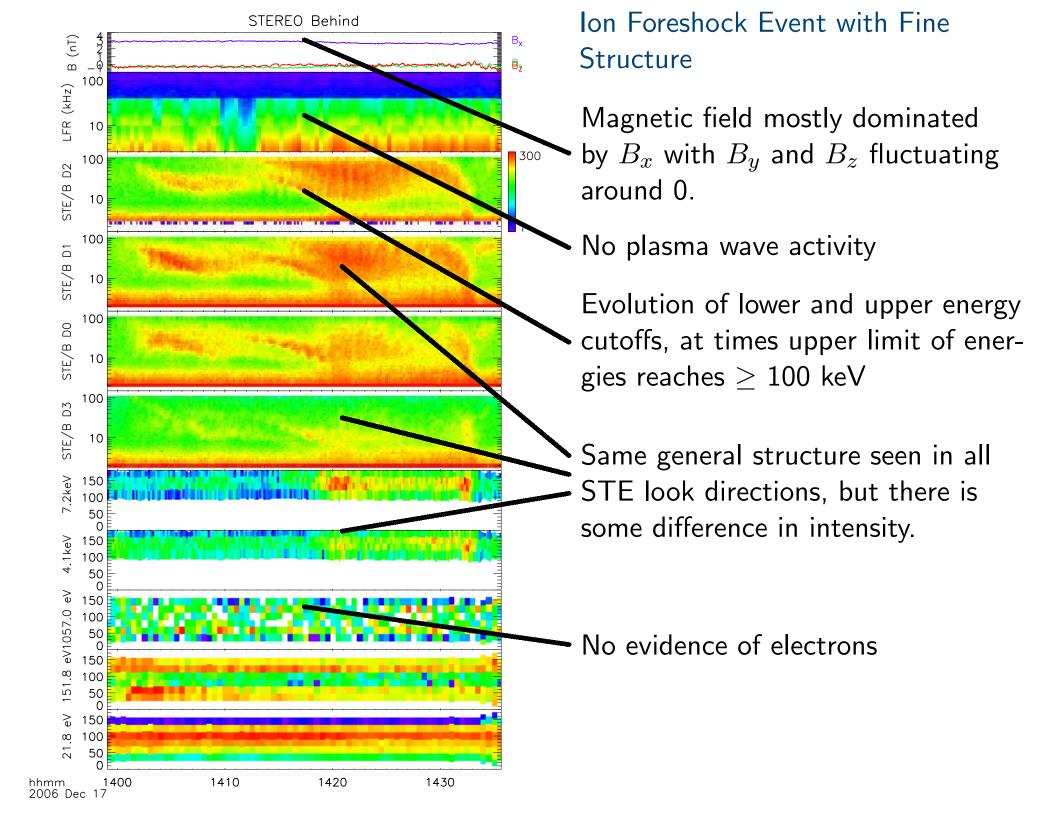
#### Ion spectra

Foreshock event with fine structure Distribution of ion events

Conclusions and Future Work



• Blue = Parallel, Red = Perpendicular.



## Distribution of ion events

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Early STEREO Orbit

Electron Foreshock

Ion Foreshock

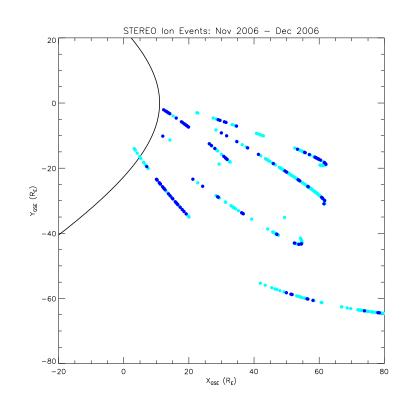
Ion Foreshock

Ion spectra

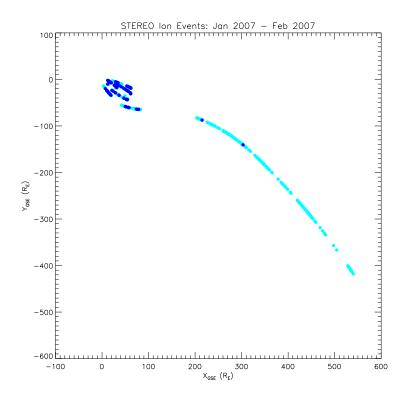
Foreshock event

with fine structure

Distribution of ion events



Light blue dot is an ion event, dark blue dot is an ion event with associated ULF activity.



Ion events continue, but ULF events die out. Issue of distance versus depth.

## **Conclusions and Future Work**

Introduction

Early STEREO Orbit

Electron Foreshock

Ion Foreshock

Conclusions and Future Work

- STE and SWEA see many examples of electrons and ions upstream of the foreshock. STEREO/A continues to observe upstream particles well into 2007.
- STE can measure the energy spectrum of foreshock electron beams with good sensitivity and energy resolution. However, it is limited by incomplete pitch angle coverage.
- STE also sees ion events, including ULF events. The strength and frequency of ULF events decreases as STEREO/A moves farther from the bow shock, but it is unclear whether this is due to distance or depth in the foreshock.
- Put observations in magnetic coordinates (organized by  $\theta_{BN}$  and depth in foreshock) rather than GSE coordinates. A better bow shock model is necessary.
- Resolve ambiguous events and use current SWEA data.

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Conclusions and Future Work

Conclusions and Future Work

# Thank you!